

# Constraints on Spacetime non-commutativity from CMB and LSS

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# Non-commutative Spacetime model

- \* Quantum theories of gravity suggest that spacetime is non-commutative at length scales close to the Planck length.

$$[\hat{x}_\mu, \hat{x}_\nu] = i\theta_{\mu,\nu}$$

- \* Introduces a special spatial direction:  $\vec{\theta}^0 = \theta\hat{\theta}^0$

# Modified Power Spectrum

\* Modified primordial scalar power spectrum:

$$P_{\theta}(\vec{k}) = P_0(k) \left[ 1 + \frac{H^2}{2} (\vec{\theta}^0 \cdot \vec{k})^2 \right]$$

\* Modified CMB power spectrum:

$$\mathcal{C}_{\ell}^{TT} = \int dk k^2 P_0(k) \left| \Delta_{\ell}^T(k) \right|^2 \frac{\sinh(\theta H k)}{\theta H k} \quad \mathcal{C}_{\ell}^{EE} = \int dk k^2 P_0(k) \left| \Delta_{\ell}^E(k) \right|^2 \frac{\sinh(\theta H k)}{\theta H k}$$

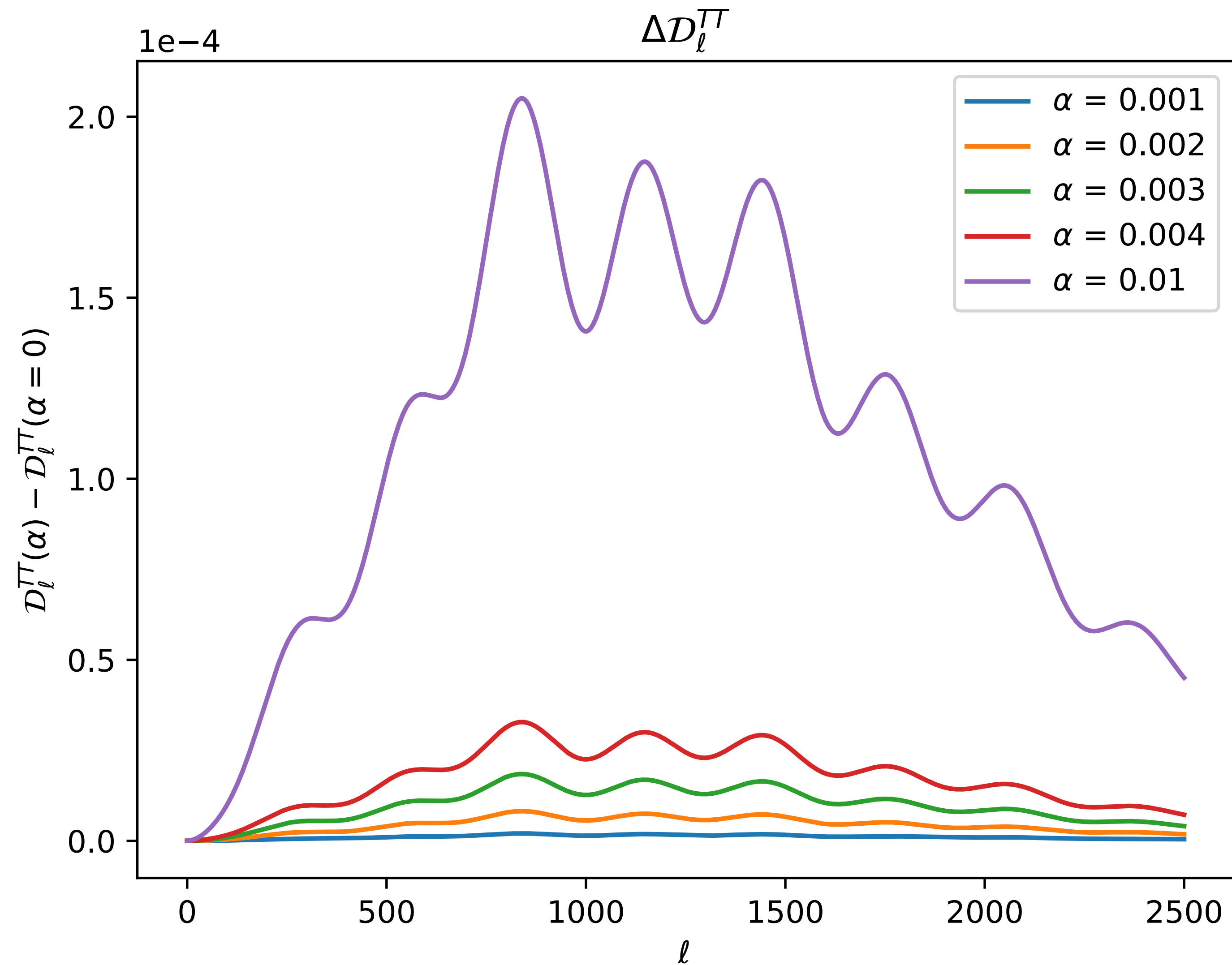
# Method

- \* Modify the primordial scalar power spectrum in CAMB to accommodate the additional  $\frac{\sinh(\theta Hk)}{\theta Hk}$  factor.
- \* Compute modified CMB power spectra for noncommutative spacetime.
- \* Run MCMC with modified CAMB to get constraints on  $\alpha = \theta H$ .
- \* Use  $H$  during inflation to constrain  $\theta$ .

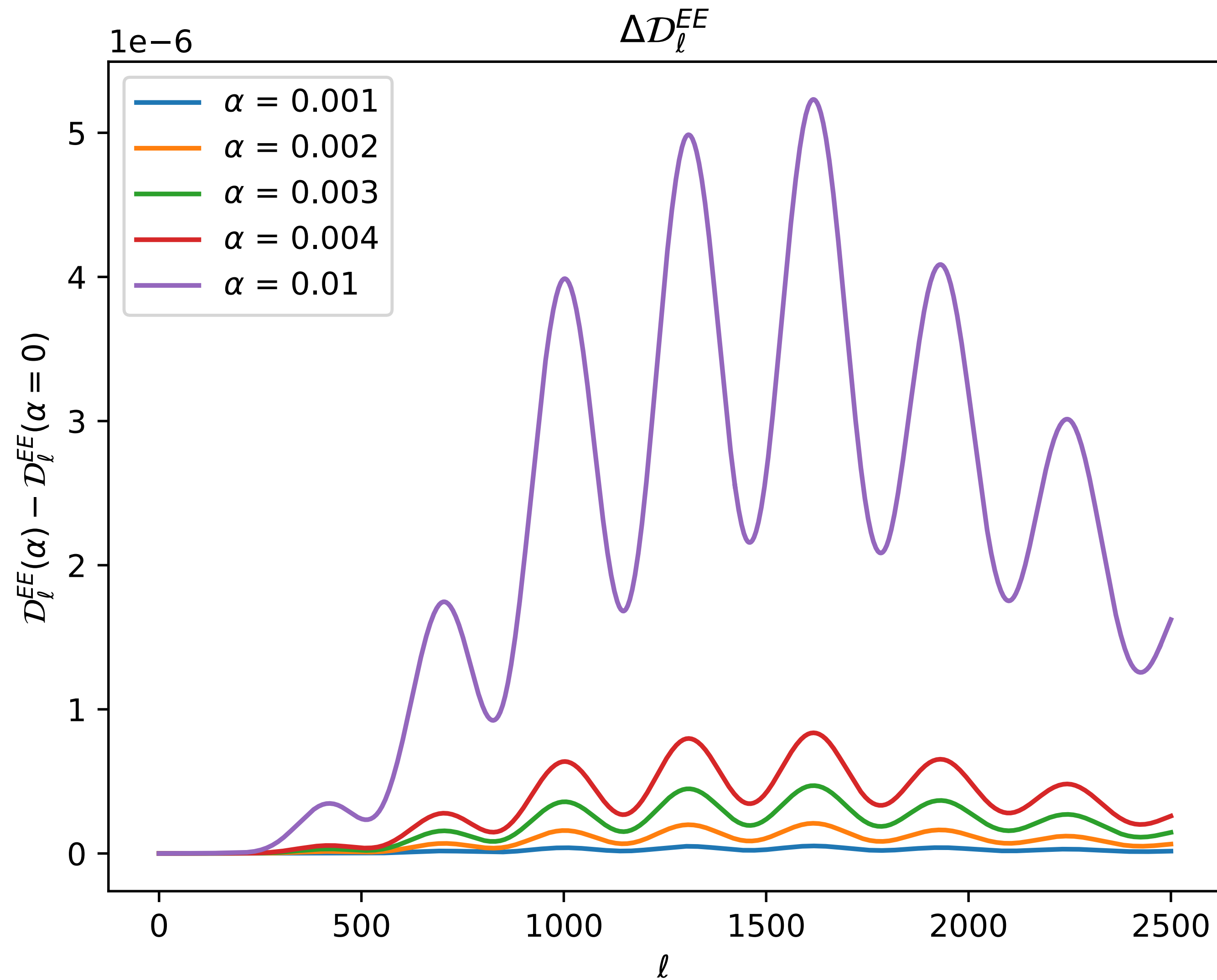
# Previous Constraints

- \* Akofor et al. (2009), with WMAP, ACBAR and CBI data got  $\sqrt{\theta} < 1.36 \times 10^{-19} \text{m}$ .
- \* Joby et al. (2015), with Planck 2013, got  $\sqrt{\theta} < 0.653 \times 10^{-19} \text{m}$ .
- \* Improvement by a factor of  $\approx 2$ , comes from higher angular resolution and precision of Planck data.

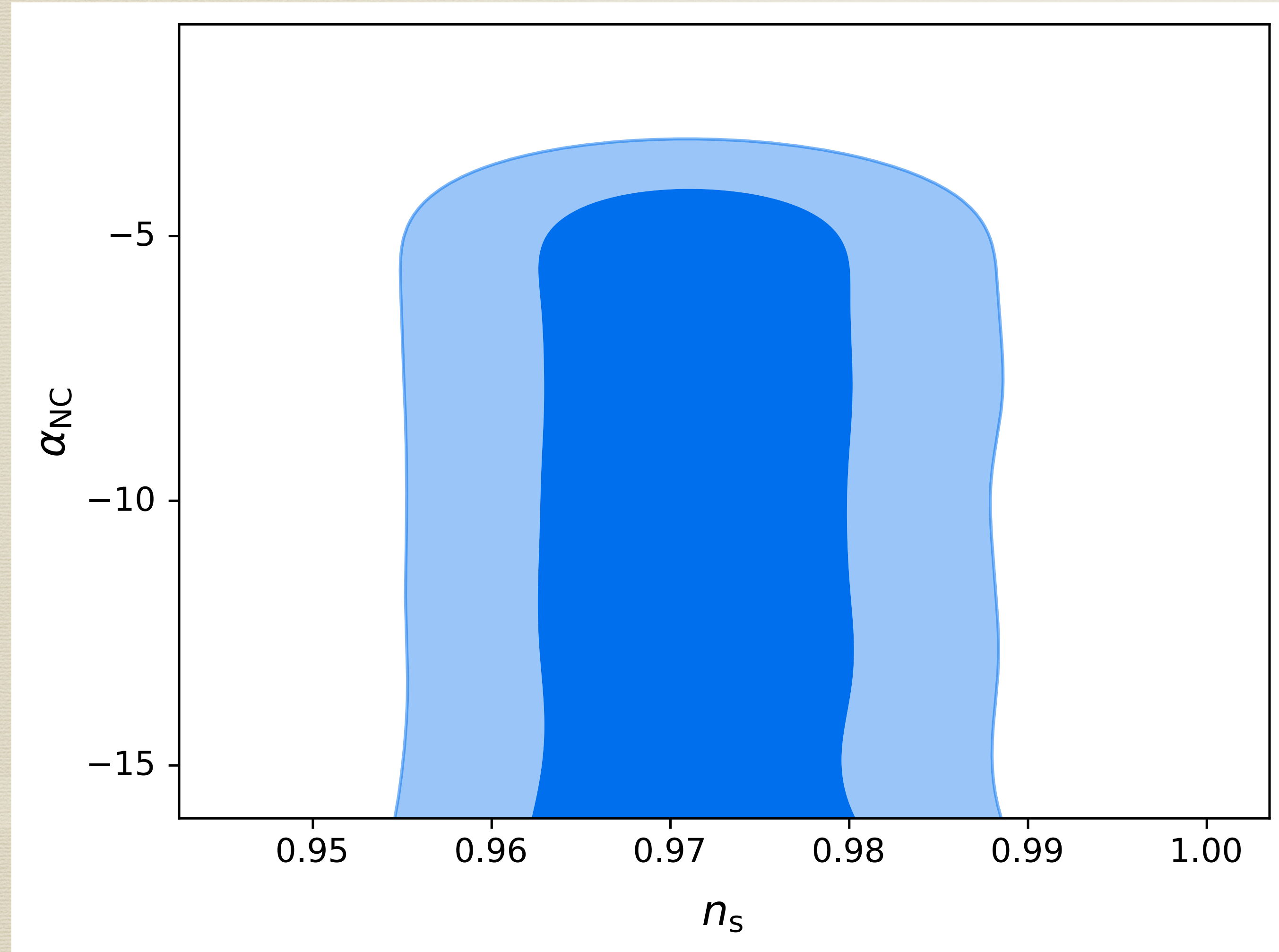
# Modified CMB (TT) power spectrum



# Modified CMB (EE) power spectrum



# Constraint from Planck 2018 data



\* Marginalised upper bound:  
 $\log_e(\alpha) < -7.66$

\*  $\alpha < 0.00047$

\* Length scale  
 $\sqrt{(\theta)} < 0.277 \times 10^{-19} \text{m}$



# How far can we get with CMB data?

- \* Fisher matrix estimate for future CMB experiments.
- \* For  $\ell_{max} > 3000$ , with error bars comparable to cosmic variance, what is the expected constraint on  $\theta$ ?
- \* At which  $\ell_{max}$  do the constraints saturate?

# Constraint from LSS simulations

- \* Linear matter power spectrum from modified CAMB for noncommutative spacetime.
- \* N-body simulations to estimate late time matter power spectrum.
- \* Constrain  $\theta$  by comparing with simulations of commutative spacetime.
- \* How does this constraint compare to the constraint from CMB data?

# Summary

- \* We have considered the prediction of noncommutative spacetime for the CMB and matter power spectra.
- \* The modification introduces a factor that increases exponentially at smaller scales.
- \* Constraint from Planck 2018 data:  $\sqrt{(\theta)} < 0.277 \times 10^{-19} \text{m}$
- \* Forecast constraint from future CMB experiments.
- \* Constraint from simulations of matter density distribution.